# Extended Waveform Tomography as Tomography

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# **Apology Haiku**

Waveform tomography...

terrible name -

really, waveform inversion



# **Principles**

- Extended modeling permits data fit, so avoids cycle skip
- Linear inversion based velocity estimation ("IVA") via differential semblance avoids gradient artifacts
- At heart, IVA is tomography





#### Hug your data

#### Data fit $\Rightarrow$ bettter gradients

The Inner Tomographer















 $x_{s=6}$ 

Born data,  $x_s = 6 \text{ km}$ 





Muted Born data,  $x_s = 6$  km





Resim from inversion, Marmousi smoothed v





Residual, Marmousi smoothed v: RMS  $\simeq$  .08





Resim from inversion,  $H_2O$  v







#### Moral: no cycle skip if you fit the data!





Hug your data

#### Data fit $\Rightarrow$ bettter gradients

The Inner Tomographer



#### "Traditional" differential semblance

$$\min_{c}(\tilde{J}_{0}[c] = \sum_{x,z,h} |hI(x,z,h)|^{2})$$

I(x, z, h) =space-shift image volume



# "Traditional" differential semblance

I(x, z, h) at c = 2.5km/s (true c = 3km/s)

ERTM with 2.5 km/s





thanks: Y. Liu

### "Traditional" differential semblance

Gradient:

$$abla ilde{J}_0[c] \simeq rac{\delta I}{\delta c}^*(h^2 I)$$

$$\frac{\delta I^*}{\delta c}$$
 = "tomographic operator"



### "Traditional" differential semblance RTM-based DS gradient, (wrong) c = 2.5km/s





thanks: Y. Liu

# Inversion VA, aka EFWI

EFWI gradient, (wrong) c = 2.5km/s





thanks: Y. Liu



Hug your data

#### Data fit $\Rightarrow$ better gradients

#### The Inner Tomographer



Extended modeling:

- *m* depends on non-physical space-time degrees of freedom
- physical models are extended: m physical  $\Leftrightarrow Am = 0$
- $\mathcal{F}$  is ordinary modeling on physical models



Extended FWI = FWI based on extended modeling, with penalty for non-physicality:

$$J[m] = \frac{1}{2} \|\mathcal{F}[m] - d\|^2 + \frac{\alpha^2}{2} \|Am\|^2$$



Example: horiz. space-shift (or subsurface offset) extended modeling, constant density acoustics in Born approximation

m = (background model c(x, z), extended reflectivity <math>r(x, z, h))

 $\mathcal{F}[m] = F[c]r$  = pressure perturbation sampled at src/rcvrs - space-shift *demigration* 



physical models =  $(c(x, z), r(x, z)\delta(h))$ 

rel'n to Born:  $r = 2\delta c/c$ 

reg op A = annihilator of physical models

Expl: multiplication by h (many other possibilities, eg. Albertin 09, Sava & Yang 12)



Born-based EFWI for acoustics:

$$J[c, r] = \frac{1}{2} \|F[c]r - d\|^2 + \frac{\alpha^2}{2} \|Ar\|^2$$

Main fact: J[c, r] is as oscillatory as data



Reduced objective:

$$\tilde{J}[c] = \min_{r} J[c, r] = J[c, r[c]]$$

where

$$r[c] = N[c]^{-1}F[c]^*d, \ N[c] = F[c]^*F[c] + \alpha^2 A^*A$$

NB: this is variable projection!



$$ilde{J}[c] = rac{1}{2} \| (F[c]N[c]^{-1}F[c]^* - I)d \|^2 + \ rac{lpha^2}{2} \| AN[c]^{-1}F[c]^*d \|^2$$

Both terms have form  $\langle d, Pd \rangle$ , P = pseudodiffl op with symbol smooth in c

 $\Rightarrow$  smooth independent of data spectrum [see Stolk & S. 03 for choice of A]



# For this problem, VPM completely changes character of objective



$$\nabla \tilde{J}[c] = DF[c]^*(r[c], F[c]r[c] - d)$$

$$DF[c]^* = D^2 \mathcal{F}[c]^* = ext{ ``tomographic operator''}$$



Key to understanding Hessian:

$$DF[c]\delta c = F[c](Q[c]\delta c)$$

where  $Q[c]\delta c$  is pseudo of order 1:

$$(Q[c]\delta c)r(z,x,h) = \int \int \int \int dk_z dk_x dk_h imes$$
  
 $(\delta \tau(x_s, z, x-h) + \delta \tau(x_r, z, x+h))(ik_z)$   
 $\hat{r}(k_z, k_z, k_h)e^{i(k_z z+k_x x+k_h h)}$ 



 $\delta \tau = D\tau[c]\delta c = \text{traveltime perturbation}$ 

 $x_s, x_r$  determined by  $z, x, h, k_x/k_z, k_h/k_z$ 

Relation generally complex, but at h = 0,

$$k_x/k_z = an\psi$$
  
 $k_h/k_z = an heta$ 

 $\psi = \mathrm{dip}$  angle,  $\theta = \mathrm{scattering}$  angle



How to see this: express  $DF[c]\delta c$  as GRT, apply pseudo inverse on left (cf. Jie's talk), use stationary phase



Compute Hessian at consistent point:

$$\begin{aligned} F[c]r &= d\\ Ar &= 0 \end{aligned}$$

Use DF[c]=F[c]Q[c], h = 0; lots of cancellations occur:

$$D^{2}\tilde{J}[c](\delta c_{1}, \delta c_{2}) = \\ \alpha^{2} \langle A(Q[c]\delta c_{1})r, (I - AN[c]^{-1}A^{*})A(Q[c]\delta c_{2})r \rangle$$



Since Ar = 0, AQr = [A, Q]r. Calc. of pseudos: symbol of commutator = Poisson braced of symbols

Symbol of 
$$A = h$$
, symbol of  
 $Q = (\delta \tau_s + \delta \tau_r)(..., \tan \psi, \tan \theta)ik_z$   
 $\Rightarrow$  symbol of  $[A, Q] = \delta p_s \frac{\partial x_s}{\partial \tan \theta} + \delta p_r \frac{\partial x_r}{\partial \tan \theta}$ 

$$p_r = rac{\partial au_r}{\partial x_r}$$
 etc



Upshot:  $D^2 \tilde{J}[c](\delta c_1, \delta c_2)$  is weighted integral of perturbations in slownesses, weighted by

- energy in reflectivity
- geometric factor (rate of change of scattering angle wrt src, rec x)

 $\Rightarrow$  near consistent data, DSO  $\simeq$  a form of slope tomography.



# Where from here

- better understanding of tomography problem at heart of space-shift DSO: linear combinations of source, receiver slope.
- ▶ better formalize role of reflectivity denser reflectors ⇒ better resolution?
- ► 3D



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